

FYS 610 Many-particle quantum mechanics

Suggested solutions, exercises for 17 February 2017

PROBLEM 6: See Suggested solutions for 10.02 2017.

PROBLEM 8: See Suggested solutions for 10.02 2017.

PROBLEM 10:

a) Making a partial integration, the lagrangian density can be rewritten:

$$\mathcal{L} = \frac{1}{2}(\partial_\mu \phi)^2 + \frac{1}{2}m^2\phi^2 - \frac{\lambda}{4!}\phi^4 = \frac{1}{2}\dot{\phi}^2 - \frac{1}{2}(\nabla\phi)^2 + \frac{1}{2}m^2\phi^2 - \frac{\lambda}{4!}\phi^4,$$

which leads to the Euler-lagrange equation:

$$\partial_\mu \frac{\partial \mathcal{L}}{\partial \partial_\mu \phi} - \frac{\partial \mathcal{L}}{\partial \phi} = \square\phi - m^2\phi + \frac{\lambda}{6}\phi^3 = 0.$$

The constant solutions, $\phi = c$, follow as:

$$\frac{\lambda}{6}c^3 - m^2c = 0 \quad \implies \quad c_0 = 0, \quad c_+ = \sqrt{\frac{6m^2}{\lambda}}, \quad c_- = -\sqrt{\frac{6m^2}{\lambda}}.$$

The energy density is equal to the Hamiltonian density:

$$\mathcal{H} = \frac{\partial \mathcal{L}}{\partial \dot{\phi}}\dot{\phi} - \mathcal{L} = \frac{1}{2}\dot{\phi}^2 + \frac{1}{2}(\nabla\phi)^2 - \frac{1}{2}m^2\phi^2 + \frac{\lambda}{4!}\phi^4,$$

which for the three constant solutions found yields the energy densities:

$$\mathcal{E} = \mathcal{T}^{00} = \mathcal{H} = -\frac{m^2c^2}{2} + \frac{\lambda c^4}{24} = 0, -\frac{3}{2}\frac{m^4}{\lambda}, -\frac{3}{2}\frac{m^4}{\lambda}.$$

Thus we see that $c_\pm = \pm\sqrt{6m^2/\lambda}$ both yields minima, which are even *global minima*, *i.e.* they both yield the ground state energy density of the system, since the non-constant terms in \mathcal{H} are positive. On the other hand, $c_0 = 0$ is a *local point of inflection* of the full \mathcal{H} , because we can decrease \mathcal{H} by adding small constant values to $\phi = c_0$, but increase it by adding space- or time-dependent values.

b) Under the symmetry transformation $\phi \rightarrow -\phi$ we see that $c_+ \leftrightarrow c_-$, so neither state is invariant.

- c) Let $c = c_+$ for definiteness. Since it is constant, we have $\partial_\mu \pi = \partial_\mu \phi$, so the Lagrangian density expressed in terms of π reads,

$$\begin{aligned}\mathcal{L} &= \frac{1}{2}\dot{\pi}^2 - \frac{1}{2}(\nabla \pi)^2 + \frac{1}{2}m^2(\pi + c)^2 - \frac{\lambda}{4!}(\pi + c)^4 \\ &= \frac{1}{2}\dot{\pi}^2 - \frac{1}{2}(\nabla \pi)^2 + \frac{1}{2}m^2\pi^2 - \frac{\lambda c}{4}\pi^3 - \frac{\lambda}{4!}\pi^4 + \frac{3}{2}\frac{m^4}{\lambda}.\end{aligned}$$

where we have substituted the value of c^2 . The equation of motions now become:

$$= \square \pi + 2m^2\pi + \frac{3\lambda c}{4}\pi^2 + \frac{\lambda}{6}\pi^3 = 0,$$

which indeed has a solution $\pi = 0$. The transformation $\phi \rightarrow -\phi$ corresponds to the combined transformation $(\pi, c) \rightarrow (-\pi, -c)$, which is a symmetry of \mathcal{L} . [We can also read off from \mathcal{L} that after quantization π gives rise to a particle of *positive* rest mass $\sqrt{2}m$.]

PROBLEM 11:

- a) We find the Euler-Lagrange equations as in *Schwartz* sec. 3.4, with an additional term in \mathcal{L} :

$$\partial_\mu F^{\mu\nu} + m^2 A^\mu = J^\nu.$$

Since $F^{\mu\nu} = -F^{\nu\mu}$, we have $\partial_\mu \partial_\nu F^{\mu\nu} = -\partial_\nu \partial_\mu F^{\nu\mu} = 0$ (this is also seen by an explicit calculation), we find if J^μ is conserved ($\partial_\mu J^\mu = 0$):

$$0 = \partial_\mu \partial_\nu F^{\mu\nu} = \partial_\nu J^\nu - m^2 \partial_\nu A^\nu = m^2 \partial_\nu A^\nu \quad \implies \quad \partial_\nu A^\nu = 0,$$

so A^μ satisfies the Lorentz condition.

- b) Like in *Schwartz* sec. 3.4, the Lorentz condition implies:

$$\partial_\mu F^{\mu\nu} = \partial_\mu \partial^\mu A^\nu - \partial_\mu \partial^\nu A^\mu = \square A^\nu$$

so the equations of motion can be written as four linear partial differential equations:

$$\square A^\mu + m^2 A^\mu = J^\mu.$$

In the present case we shall find the static potential from a point charge at the origin, like in *Schwartz* sec. 3.4.2, so $A^\mu = (\rho, \mathbf{0})$, with $\rho = e\delta^3(\mathbf{x})$. Hence $\mathbf{A} = \mathbf{0}$, while A^0 is a solution of:

$$(-\nabla^2 + m^2) A^0(\mathbf{x}) = e\delta^3(\mathbf{x})$$

Taking Fourier transforms, this becomes:

$$(\mathbf{k}^2 + m^2) \tilde{A}^0(\mathbf{k}) = e \quad \implies \quad A^0(\mathbf{k}) = \frac{e}{k^2 + m^2},$$

where $k = |\mathbf{k}|$. Transforming back, we find by introducing spherical coordinates, like in *Schwartz* eq. (3.62):

$$\begin{aligned}A^0(\mathbf{x}) &= e \int \frac{d^3\mathbf{k}}{(2\pi)^3} \frac{e^{i\mathbf{k}\mathbf{x}}}{\mathbf{k}^2 + m^2} = \frac{e}{4\pi^2} 2\pi \int_0^\infty \frac{k^2 dk}{k^2 + m^2} \frac{e^{ikr} - e^{-ikr}}{ikr} \\ &= \frac{e}{4\pi^2 r i} \int_{-\infty}^\infty \frac{k dk}{k^2 + m^2} e^{ikr}.\end{aligned}$$

- c) This integral can be evaluated by contour integration (or be looked up in an integral table). Since $r > 0$, we must close the contour in the upper half plane. Since

$$\frac{k}{k^2 + m^2} = \frac{k}{(k + im)(k - im)},$$

we see that there is only one pole, at $k = im$, in this plane, with residue:

$$\text{Res}_{k=im} = \frac{im}{2im} e^{-mr} = \frac{1}{2} e^{-mr}.$$

(this is the same as in *Schwartz*, except $\delta \rightarrow m$, which remains finite). Hence we have:

$$A^0(\mathbf{x}) = \frac{e}{4\pi^2 r i} 2\pi i \text{Res}_{k=im} = \frac{e}{4\pi} \frac{e^{-mr}}{r}$$

- d) This follows trivially, since $\lim_{m \rightarrow 0} e^{-mr} = 1$.
e) Introducing standard units, the Yukawa potential is:

$$A^0(r) = \frac{e}{4\pi\epsilon_0} \frac{e^{-cmr/\hbar}}{r} = \frac{e}{4\pi\epsilon_0} \frac{e^{-r/r_s}}{r}.$$

Here $r_s = \hbar/mc$ is a screening length, while $2\pi r_s$ is called the *Compton wavelength*. For distances $r \ll r_s$, the Yukawa and the Coulomb potentials are the same, but for $r \gg r_s$ the former vanishes, so it only has a finite range. It was introduced to describe the short range of nuclear forces, so r_s was thought to be the range of the strong nuclear force, around 10^{-15}m . Yukawa predicted that there should exist particles, now identified with the pion, with a corresponding rest energy $mc^2 = \hbar c/r_s \approx 200\text{MeV}$. Indeed, the pion mass is 139.5MeV . This screened potential had actually been introduced in physics earlier by Peter Debye, to describe electric screening in plasmas and electrolytes, for which it is still in wide use.

- e) We can rewrite the free, massless part of \mathcal{L} , interchanging indices $\mu \leftrightarrow \nu$ and making further index gymnastics necessary to identify equal terms:

$$\begin{aligned} \mathcal{L}_0 &= \frac{1}{4} F^{\mu\nu} F_{\mu\nu} = \frac{1}{4} (\partial^\mu A^\nu - \partial^\nu A^\mu) (\partial_\mu A_\nu - \partial_\nu A_\mu) \\ &= \frac{1}{2} \left((\partial^\mu A^\nu)^2 - (\partial^\nu A^\mu) (\partial_\mu A_\nu) \right) \\ &= \frac{1}{2} \left((\partial^\mu A^\nu)^2 - \partial^\nu [A^\mu (\partial_\mu A_\nu)] + A^\mu \partial_\mu \partial^\nu A_\nu \right) \\ &= \frac{1}{2} \left((\partial^\mu A^\nu)^2 - \partial^\nu [A^\mu (\partial_\mu A_\nu)] + \partial_\mu [A^\mu \partial^\nu A_\nu] - (\partial_\mu A^\mu) (\partial^\nu A_\nu) \right) \\ &= \mathcal{L}'_0 + \partial_\mu X^\mu. \end{aligned}$$

Here

$$\mathcal{L}'_0 = \frac{1}{2} \left((\partial^\mu A^\nu)^2 - (\partial_\mu A^\mu) (\partial_\nu A^\nu) \right)$$

only differs from \mathcal{L}_0 by a the total derivative $\partial_\mu X^\mu$, with $X^\mu = A^\mu \partial^\nu A_\nu - A^\nu \partial^\mu A_\mu$, and so gives rise to the same equations of motion. If we plug $\partial_\mu A^\mu = 0$ into \mathcal{L}'_0 , we simply have $\mathcal{L}'_0 = \frac{1}{2} (\partial^\mu A^\nu)^2$, which also has Euler-Lagrange equation $\square A^\mu = 0$. Thus the equations of motion are unchanged when the constraint is included from the start. It is the term $\frac{1}{2} m^2 A_\mu^2$ in the original Lagrangian which enforces it, but, of course, without it we have a theory that differs in more respect than just the absence of the constraint.